Impact of Sunshine Duration and Clearness Index on Diffuse Solar Radiation Estimation in Mountainous Climate

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In this paper, measured data of solar radiation was applied to develop forty-three (43) empirical models for estimation of monthly average diffuse solar radiation using clearness index, sunshine duration and a combination of them as predictors. The data covered a period of two years from May 2015 to April 2017 and was measured at Mehran University of Engineering and Technology, Hyderabad, Pakistan. Through a comprehensive statistical performance analysis, 43 dimensional models developed were tested for constructing the most accurate regression model to predict the monthly mean daily diffuse solar radiation in Hyderabad, Pakistan. On the whole, the model 42 - a hybrid of sunshine duration and clearness index predictors of diffuse fraction outperformed the remaining models proposed in this study. The best model (model 42) was then compared with 5 models and 5 measured data of diffuse solar radiation available in the literature and the NASA database by applying statistical indicators such as MBE, MPE, RMSE, RRMSE, R2 and GPI. Through the analysis, the hybrid of sunshine duration and clearness index predictors of diffuse fraction model (model 42) was selected as the most appropriate model. The study concluded that the proposed hybrid model can serve as a baseline for the design of photovoltaic systems and estimate the monthly mean daily diffuse solar radiation on the horizontal surface for Hyderabad, Pakistan and other locations with similar local climate conditions.

Keywords: Diffuse solar radiation; Empirical models; Classification; Functional forms; Hyderabad; Pakistan

1. Introduction

Since the beginning of the 19th century, the exploitation of conventional fuels is increasingly moving towards the development of industrialization and modern life style. It has resulted in various health hazards, environmental pollution, disruption of ecosystems such as crop and animal diversity, increased global warming and many more factors which drive the earth towards a dark future. Thus, the world needs a smart energy source that is unlimited in reserve and can be applied without major contributions to atmospheric pollution and greenhouse effect.

As reported by the literature [1], the earth has been already presenting numerous signs of global climate change as follows. NASA Goddard Institute for space studies

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reported that the average temperature has climbed 0.8 degrees Celsius across the globe since 1880 and much of this in recent decades. The UN Intergovernmental Panel on Climate Change (IPCC) recorded that the last two decades of the 20th century have been the hottest in 400 years and possibly the warmest for several millennia [2]. The multinational Arctic Climate Impact Assessment reported that the Arctic is sensitive to atmospheric pollution and greenhouse effect. According to a report compiled between 2000 and 2004, the average temperatures in Alaska, Western Canada, and Eastern Russia have twice the global average. According to the IPCC 2007 report, based on the work of about 2,500 scientists in more than 130 countries, humans have caused all or most of the current planetary warming often called anthropogenic climate change. Industrialization, deforestation, and pollution have greatly increased atmospheric concentration of water vapor, carbon dioxide, methane, nitrous oxide, and all greenhouse gases that help trap heat near the Earth's surface. In fact, some experts worry that the natural cycles in Earth's orbits can alter the planet's exposure to sunlight, which may explain the current trend. Earth has indeed experienced warming and cooling cycles roughly every hundred thousand years due to these orbital shifts, but such changes have occurred over the span of several centuries. Nowadays change has taken place over the past 100 years or less.

It is therefore imperative for human beings to set a different course in its need for energy. For example, those involve less intrusive sources such as renewable energy sources, which do not harm the planet but still are inexhaustible.

In order to correct these anomalies, researchers, scientists, governmental and non-governmental organizations are striving tirelessly on renewable energy, which should be commercially viable, pollutant free, easy to access, and must be widespread in nature [3]. As a result of its minimal or zero impact on the environment and ecosystem which eventually poses no health hazards to man and animals, renewable energies such as solar energy, wind energy, hydro power, tidal energy and biofuels are more suitable compared to fossil fuel sources of energy, as these are generated from natural processes such as sunlight, wind, rain and numerous forms of biomass [4]. These smart energy sources are not only renewable but also have the ability to sustain ecology and the environment, as they are eco-friendly and do not contribute to global warming and production of greenhouse gases [3].

Solar energy among other renewable energy types has remained the most viable source, which has the capacity to sustain and maintain all activities and processes, support life of animals, heat the atmosphere and lands, generate wind, drive the water cycle, warm the ocean, generate fossil fuels, and support life of plants [3].

Solar energy is not only primarily derived from solar radiation reaching the earth's surface, but it remains important in many industrial and application areas, such as exciting electrons in a photovoltaic cell, solar heating, solar architecture, molten salt power plants and supplying energy to natural processes [3]. It also plays a major role in de-carbonizing the global economy and improving costs of greenhouse gas emitter [5]. Much more, solar radiation level is equally used to determine the type of photovoltaic technology (such as network-connected systems (on-grid), network-connected home systems, network-connected solar power plants, off-grid, hybrid systems and independent systems for economic purposes), concentrated solar power technology (like dish stirring, parabolic trough, linear Fresnel or central tower), and photovoltaic thermal (PVT, such as air-based PVT, water-based PVT, bifluid based PVT, PVT nanofluid based, PVT based phase change materials, PVT refrigerant based, PVT heat pipe based and PVT with heat pump) that should be installed at a particular site. It could also serve as a baseline for

estimating and understanding the diurnal fluctuations in multiple solar radiation parameters such as direct normal irradiance, diffuse horizontal irradiation, ground reflected solar radiation, evaporation and reference evapotranspiration.

In fact, despite the influence of the government, scientists, researchers and investors have explored solar energy as a type of renewable energy using the abovementioned various technologies, but fundamentally, the potential of solar energy has not been fully utilized [3]. For example, the energy emitted by the sun is so high that when only 0.1% of solar energy reaching the ground is converted into electricity with only 10% efficiency, the power output will be 17,300 GW, which is 7 of the global average instantaneous power consumption in 2012 [6-8].

These are significant and potential solar radiation reaching the Earth's surface in the form of solar energy, measurement of solar radiation and its components such as diffuse solar radiation and direct normal irradiance is limited, because there are very few standard weather stations can measure. Routinely, the data for these parameters is not available in the site of interest. However, other meteorological and atmospheric variables such as ambient temperature, cloud cover, rainfall regimes, and relative humidity are often measured routinely in most weather stations across the globe, as a result of its direct application in agricultural sciences and meteorology.

Due to the cost implication, maintenance, expertise needed for ground and satellite-derived technique of measuring solar radiation data (especially in rural and developing countries), prediction of solar radiation over a particular location using mathematical models has been initiated by solar energy researchers. Mathematical modeling serves as an alternative technique of generating data of solar radiation and its components without instrumentation network that would otherwise be needed.

Some researchers have stressed that accurate determination of diffuse solar radiation is important in design and performance analysis of solar energy projects, such as for designing and sizing photovoltaic sources as the future alternative energy [9-11]. For instant, Khorasanizadeh *et al.* [12] revealed that the impact of diffuse solar radiation to the annual solar energy is nearly 20% in Tabass, Iran.

It has been observed that in different locations across the globe, ground measurement of diffuse solar radiation is either scare or absent, whereas ground measurement of global solar radiation and weather parameters such as sunlight hours and precipitation are often available as a result of their traditional use in building and construction industries, agriculture, and meteorology. By applying mathematical correlations, diffuse solar radiation can be obtained as far as global solar radiation and other popularly measured meteorological parameters are available. For this reason, solar energy researchers across the globe have developed numerous empirical models in most metropolitan cities, because most meteorological and weather stations are often situated in these locations. From the mid-19th century, solar energy researchers have developed various empirical models for estimating diffuse solar radiation employing popularly measured variables. These variables include minimum and maximum temperature, hours of solar radiation and relative humidity [13-18]. Several researchers have equally developed regression models for estimating the monthly mean daily diffuse solar radiation employing the clearness index [4, 6, 9,13, 19-31] or applying sunshine hour fraction [28, 30, 32-36] or with combination. of them [18, 37-41]. Despotovic et al. [42] observed that the empirical models using both clearness index and sunshine duration offer better estimation of diffuse solar radiation in the main five climate zones according to Koppen-Geiger climate classification

In spite of the vast number of studies on empirical models for estimating solar radiation across the globe, there is no recorded study in Hyderabad, Pakistan. The main objective of this study was to estimate forty-three models employed for estimating diffuse solar radiation using sunshine duration, clearness index and both of the predictors, obtain the best performing model using statistical indicators (such as mean bias error (MBE), mean percentage error (MPE), root mean square error (RRMSE), coefficient of determination (R²) and global performance indicator (GPI), and compare the selected best models with five models developed from the literature and five ground measured diffuse solar radiation in the literature together with satellite data obtained from NASA database for estimating diffuse solar radiation in Hyderabad, Pakistan.

2. Materials and Methods

2.1 Study Area

Hyderabad lying along the Indus River is the second largest city of Sindh province and the 8th largest city in Pakistan (Fig. 1). It has a relatively mountainous climate which is slightly more pleasant than other parts of Central Sindh throughout the year. Summer and winter are the two main seasons, while spring and autumn are very short. The period from mid-April to late June is the hottest time of the year with temperatures as high as 48.5 °C. Winters are usually warm, around 25 °C during the day time and often below 10 °C at night, and are the best time to visit the city. The highest ever recorded temperature in Hyderabad was 48.5 °C in 1991, while the lowest was 1 °C in 2012 [43, 44].



Fig. 1. Map of Pakistan showing the study site. Hyderabad at the southern part of the figure [43].

2.2 Acquisition of Data

Under the present study, the measured global solar radiation and its components (diffuse and direct solar radiation) together with other meteorological parameters were measured by the Energy Sector Management Assistance Program of The World Bank Group at Mehran University of Engineering and Technology (M-UET), Hyderabad, Pakistan [45]. The measurements were performed for a period of two years (May, 2015 -April, 2017) so as to determine regional solar radiation and other meteorological variables. However, the monthly mean daily sunshine hours (for sunshine fraction in Equation 4) were based on the 30-year period (1981-1990) using the same geographical information as M-UET were obtained from the International Water Management Institute (IWWI) website [46]. The characteristics and specification of solar and other meteorological parameter instrument used are provided in Table 1. The obtained raw data (10 minutes summarization interval values) were post-processed in order to obtain daily values of global, diffuse and direct solar radiation data and other meteorological parameters such as air temperature, relative humidity, wind speed and wind direction. The data obtained was further averaged for a month so as to calculate the monthly mean values. The measured monthly mean daily values thus obtained and sunshine hours are shown in Table 2.

Table 1. Characteristics of the used solar and meteorological instruments [45].

Sensor manufacturer	Height orientation	Sensor type		Model	Serial number
CSPS	2m	RSI		Twin-RSI	CSPS.ms.14.001-
					0015
Kipp & Zonen	2m	Pyranometer		CMP10	140572
Campbell Scientific	2m	Temperature	and	CS215	E12267
		Humidity			
Setra	2m	Barometric pressure		Setra 278	6015966
NRG	10m	Anemometer		NRG40C	1795-00229369
NRG	10m	Wind direction		NRG200	

Table 2. Monthly and yearly averages of measured meteorological data for Hyderabad, Pakistan

	2. Monthly	and yeariy a	verages of m	easured met	eorological (iata for Hyd	erabad, Paki	stan	
Month/Mean	H	Hb	Hd	RH	T	WS	WD	AP	S
January	14.33	14.75	6.18	60.58	17.59	3.35	153.62	1009.97	8.88
February	18.96	20.78	6.32	36.58	21.56	3.34	168.53	1008.45	9.27
March	22.02	19.48	8.36	40.93	26.92	4.02	185.86	1003.96	9.48
April	24.84	19.00	10.29	39.56	31.33	5.28	202.47	999.12	9.91
May	26.07	19.14	10.84	50.85	33.42	7.00	211.71	995.46	10.51
June	24.78	15.90	11.97	54.57	33.67	6.29	197.65	992.00	9.45
July	22.22	10.72	13.28	65.46	31.75	7.71	213.75	991.05	7.43
August	21.98	13.39	11.26	67.25	30.56	6.52	204.56	993.60	7.67
September	22.49	20.28	8.11	61.58	30.28	5.87	200.81	997.89	9.35
October	19.24	19.91	6.40	54.38	29.46	4.11	184.60	1002.79	9.89
November	16.03	18.11	5.76	42.13	24.40	3.09	166.63	1007.28	9.37
December	14.71	18.71	5.02	47.63	20.24	2.84	166.74	1009.29	8.90
Monthly Mean	20.64	17.51	8.65	51.79	27.60	4.95	188.08	1000.89	9.18
Annual Mean	7503.759	6359.958	3153.287	18590.04	10125.66	1792.051	68239.03	366458.5	-

Where monthly values and mean of global (H), diffuse (H_d) and direct (H_b) solar radiation are in MJm⁻²day⁻¹ between May, 2015 – April 2017, annual mean of global (H), diffuse (H_d) and direct (H_b) solar radiation are in MJm⁻²year⁻¹ in 2016 only, RH represents relative humidity in (%), T stands for air temperature in (°C), WS designates wind speed in (m/s), WD represents wind direction in (0 – 360), AP stands for Ambient air pressure in (hPa) and S represents monthly mean sunshine hours downloaded from IWMI website [46] in hrs.

2.3 Basic Parameters

The fundamental requirements such as maximum possible sunshine hours (S_o) and extraterrestrial solar radiation on the horizontal surface (H_o) are significant for the prediction of diffuse solar radiation expressed as mathematically as given by Yaniktepe and Gene [47]:

$$H_O = \frac{24}{\pi} I_{SC} \left(1 + 0.033 \cos \frac{360n}{365} \right) \times \left(\cos \phi \cos \delta \sin \omega_S + \frac{2\pi \omega_S}{360} \sin \phi \sin \delta \right)$$
 (1)

Where the mean sunrise hour angle (ω_s) can be evaluated as:

$$\omega_s = \cos^{-1} \left[-\tan \delta \tan \phi \right] \tag{2}$$

The solar declination (δ) is given by Yaniktepe and as Genc [47] as:

$$\delta = 23.45 \sin \left[\frac{360(n+284)}{365} \right] \tag{3}$$

where I_{SC} is the solar constant, ϕ is the latitude and n the number of days of the year starting from first January.

The maximum possible sunshine duration is calculated as:

$$S_o = \frac{2}{15} \cos^{-1} \left[-\tan \delta \tan \phi \right] \tag{4}$$

where other symbols retain their usual meaning.

2.4 Statistical Modeling

Estimation of diffuse component of global solar radiation involves modeling the monthly mean diffuse fraction or diffuse coefficient as a function of monthly mean sunshine fraction, clearness index and combination of sunshine fraction and clearness index. This could be attributed to the fact that lower fluctuations are often observed in monthly mean values of solar radiation from one month to another as component to daily values of solar radiation [22]. Hence, better estimation capacity is observed in monthly mean models [22].

Peers and researchers have stressed that validating training dataset using the same dataset of training might lead to partially validated results [21, 48], thus, an independent validation dataset which involves that validating patterns have not been previously applied for training dataset is often employed. However, as a result of the short-term measure employed in this study (2 years), the present study employed dataset during May, 2015 - April, 2016 to develop the models for the station while validation dataset during May, 2016 - April, 2017 was used to test the models. This measure was employed to prevent the models from over fitting and to determine the estimation capacity of the developed models.

In diffuse solar radiation estimation, an empirical model uses diffuse fraction (H_d/H) or diffuse coefficient (H_d/H_o) with other easily measurable parameters. Moreover, since the first primitive work of Liu and Jordan [49] that estimated the mean diffuse solar radiation, numerous solar energy researchers have proposed several models in order to elaborate the Liu and Jordan model's functional form. The relationships representing the diffuse radiation are classified into three main classes: (1) sunshine duration-based models, (2) clearness index-based models, and (3) sunshine and clearness index-based models [50]. Owing to these classifications, the diffuse fraction (H_d/H) , and diffuse coefficient (H_d/H_o) correlations were used in estimating the diffuse solar radiation in Hyderabad, Pakistan.

2.4.1 Clearness Index-Based Models

According to Nwokolo and Ogbulieze [50], models of the monthly mean diffuse fraction (H_d/H) and the diffuse coefficient (H_d/H_o) is a function of the clearness index; such that

$$\left(\frac{H_d}{H}, \frac{H_d}{H_o}\right) \approx f\left(\frac{H}{H_o}\right) \tag{5}$$

The proposed models under this class is shown in Table 3a.

2.4.2 Sunshine Duration-Based Models

Numerous models have been widely applied sunshine fraction (S/S_o) in associating the ratio of diffuse solar radiation (H_d) to often expressed as diffuse fraction (H_d/H) , and the monthly average diffuse coefficient (H_d/H_o) to sunshine fraction or combination of both. Varying degrees of polynomial functions such as linear and quadratic, logarithmic and exponential models are applied for this study. Where S is the monthly mean daily hours of sunshine; such that

$$\left(\frac{H_d}{H}, \frac{H_d}{H_o}\right) \approx f\left(\frac{S}{S_o}\right) \tag{6}$$

The developed models under this class is shown in Table 3b.

2.4.3 Sunshine Duration and Clearness Index-Based Models

Under this class, the monthly mean diffuse fraction (H_d/H) and the diffuse coefficient (H_d/H_o) are function of the clearness index and sunshine fraction; such that

$$\left(\frac{H_d}{H}, \frac{H_d}{H_o}\right) \approx f\left(\frac{H}{H_o}, \frac{S}{S_o}\right) \tag{7}$$

The developed models under this class is shown in Table 3c.

2.5 Comparison of Models

In order to check the capacity and accuracy of the estimated data from the measured data in this study, numerous statistical indicators are applied [22-24, 42]. These metrics include mean bias error (MBE), mean percentage error (MPE), root mean square error (RMSE), relative root mean square error (RRMSE) and coefficient of determination (R²) as presented in equation (8 - 12).

$$MBE = \frac{1}{n} \sum_{i=1}^{n} \left(H_d^{est} - H_d^{mea} \right)$$
 (8)

$$MPE = \frac{100}{n} \sum_{i=1}^{n} \left(\frac{H_d^{mea} - H_d^{est}}{H_d^{mea}} \right)$$

$$\tag{9}$$

$$RMSE = \left[\frac{1}{n}\sum_{i=1}^{n} \left(H_d^{est} - H_d^{mea}\right)^2\right]^{\frac{1}{2}}$$

$$\tag{10}$$

Table 3a. Monthly mean diffuse solar radiation models under clearness index-based models

S/N	Parameters	Type	ar radiation models under clearness index-based models Regression relations	Model
1	$\frac{H}{H_o}$	Linear	$\frac{H_d}{H} = 1.646 - 1.957 \left(\frac{H}{H_O}\right)$	1
2	$\frac{H}{H_O}$	Quadratic	$\frac{H_d}{H} = 6.062 - 16.466 \left(\frac{H}{H_o}\right) + 11.8676 \left(\frac{H}{H_o}\right)^2$	2
3	$\frac{H}{H_O}$	Polynomial	$\frac{H_d}{H} = 4.581 - 9.196 \left(\frac{H}{H_o}\right) + 6.442 \left(\frac{H}{H_o}\right)^3$	3
4	$\frac{H}{H_O}$	Polynomial	$\frac{H_d}{H} = 3.841 - 6.773 \left(\frac{H}{H_o}\right) + 5.235 \left(\frac{H}{H_o}\right)^4$	4
5	$\frac{H}{H_o}$	Linear	$\frac{H_d}{H_o} = 0.721 - 0.735 \left(\frac{H}{H_o}\right)$	5
6	$\frac{H}{H_O}$	Quadratic	$\frac{H_d}{H_o} = 2.680 - 7.173 \left(\frac{H}{H_o}\right) + 5.265 \left(\frac{H}{H_o}\right)^2$	6
7	$\frac{H}{H_O}$	Polynomial	$\frac{H_d}{H_o} = 2.025 - 3.952 \left(\frac{H}{H_o}\right) + 2.862 \left(\frac{H}{H_o}\right)^3$	7
8	$\frac{H}{H_O}$	Polynomial	$\frac{H_d}{H_o} = 1.698 - 2.878 \left(\frac{H}{H_o}\right) + 2.329 \left(\frac{H}{H_o}\right)^4$	8
9	$In \frac{H}{H_O}$	Logarithmic	$\frac{H_d}{H_o} = 0.049 - 0.451 \left(In \frac{H}{H_o} \right)$	9
10	$In \frac{H}{H_O}$	Logarithmic	$\frac{H_d}{H} = -0.143 - 1.199 \left(In \frac{H}{H_o} \right)$	10
11	$\exp \frac{H}{H_O}$	Exponential	$\frac{H_d}{H} = 2.401 - 1.058 \left(\exp \frac{H}{H_o} \right)$	11
12	$\exp \frac{H}{H_O}$	Exponential	$\frac{H_d}{H} = 4.912 - 8.495 \left(\exp \frac{H}{H_o} \right) + 7.690 \left(\exp \frac{H}{H_o} \right)^2$	12
13	$\exp \frac{H}{H_O}$	Exponential	$\frac{H_d}{H_o} = 1.004 - 0.397 \left(\exp \frac{H}{H_o} \right)$	13
14	$\exp \frac{H}{H_O}$	Exponential	$\frac{H_d}{H_o} = 2.109 - 3.668 \left(\exp \frac{H}{H_o} \right) + 3.382 \left(\exp \frac{H}{H_o} \right)^2$	14

Table 3b. Monthly mean diffuse solar radiation models unde	ler sunshine duration-based models	3
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S/N	Parameters	Type	liation models under sunshine duration-based models Regression relations	Model
15	$\frac{s}{s_o}$	Linear	$\frac{H_d}{H} = 0.973 - 0.733 \left(\frac{S}{S_o} \right)$	15
16	$\frac{s}{s_o}$	Quadratic	$\frac{H_d}{H} = 1.394 - 1.952 \left(\frac{S}{S_o}\right) + 0.860 \left(\frac{S}{S_o}\right)^2$	16
17	$\frac{s}{s_o}$	Polynomial	$\frac{H_d}{H} = 1.246 - 1.328 \left(\frac{S}{S_o}\right) + 0.390 \left(\frac{S}{S_o}\right)^3$	17
18	$\frac{S}{S_o}$	Polynomial	$\frac{H_d}{H} = 1.172 - 1.122 \left(\frac{S}{S_o}\right) + 0.263 \left(\frac{S}{S_o}\right)^4$	18
19	$\frac{S}{S_o}$	Linear	$\frac{H_d}{H_o} = 0.500 - 0.317 \left(\frac{S}{S_o} \right)$	19
20	$\frac{s}{s_o}$	Quadratic	$\frac{H_d}{H_o} = 0.274 + 0.336 \left(\frac{S}{S_o}\right) - 0.461 \left(\frac{S}{S_o}\right)^2$	20
21	$\frac{s}{s_o}$	Polynomial	$\frac{H_d}{H_o} = 0.348 + 0.015 \left(\frac{S}{S_o}\right) - 0.217 \left(\frac{S}{S_o}\right)^3$	21
22	$\frac{s}{s_o}$		$\frac{H_d}{H_o} = 0.385 - 0.092 \left(\frac{S}{S_o}\right) - 0.152 \left(\frac{S}{S_o}\right)^4$	22
23	$In \frac{S}{S_O}$	Logarithmic	$\frac{H_d}{H_o} = 0.197 - 0.219 \left(In \frac{S}{S_o} \right)$	23
24	$\exp \frac{S}{S_o}$	Exponential	$\frac{H_d}{H_o} = 0.594 - 0.156 \left(\exp \frac{S}{S_o} \right)$	24
25	$\exp \frac{S}{S_o}$	Exponential	$\frac{H_d}{H_o} = 0.493 - 0.010 \left(\exp \frac{S}{S_o} \right) - 0.140 \left(\exp \frac{S}{S_o} \right)^2$	25
26	In S_o , exp S_o	Exponential and logarithmic	$\frac{H_d}{H_o} = 0.897 + 0.170 \left(In \frac{S}{S_o} \right) - 0.275 \left(\exp \frac{S}{S_o} \right)$	26
27	$In \frac{S}{S_O}$	logarithmic	$\frac{H_d}{H} = 0.270 - 0.514 \left(\ln \frac{S}{S_o} \right)$	27
28	$In \frac{S}{S_O}$	logarithmic	$\frac{H_d}{H} = 0.270 - 0.257 \left(\ln \frac{S}{S_o} \right)^2$	28
29	$\exp \frac{S}{S_o}$	Exponential	$\frac{H_d}{H} = 1.184 - 0.357 \left(\exp \frac{S}{S_o} \right)$	29

30
$$\exp \frac{S}{S_o}$$
 Exponential $\frac{H_d}{H} = 1.500 - 0.971 \left(\exp \frac{S}{S_o} \right) + 0.519 \left(\exp \frac{S}{S_o} \right)^2$ 30

31
$$In \frac{S}{S_o}$$
, exp $\frac{S}{S_o}$ Exponential and $\frac{H_d}{In S_o} = -0.114 + 0.150 \left(\exp \frac{S}{S_o} \right) - 0.727 \left(In \frac{S}{S_o} \right)$ 31

Table	e 3c. Monthly me	an diffuse solar radiation models under sunshine duration and clearness index-based n	nodels
S/N	Parameters	Regression relations	Model
32	$\frac{H}{H_o}, \frac{S}{S_o}$	$\frac{H_d}{H} = 1.368 - 0.942 \left(\frac{H}{H_o}\right) - 0.4732 \left(\frac{S}{S_o}\right)$	32
33	$\frac{H}{H_o}, \frac{S}{S_o}$	$\frac{H_d}{H} = 0.807 + 0.903 \left(\frac{H}{H_o}\right) - 0.486 \left(\frac{S}{S_o}\right) - 1.487 \left(\frac{H}{H_o}\right)^2$	33
34	$\frac{H}{H_o}$, $\frac{S}{S_o}$	$\frac{H_d}{H} = 3.638 - 11.949 \left(\frac{H}{H_o}\right) + 3.038 \left(\frac{S}{S_o}\right) + 8.437 \left(\frac{H}{H_o}\right)^2 - 2.322 \left(\frac{S}{S_o}\right)^2$	34
35	$\frac{H}{H_o}$, $\frac{S}{S_o}$	$\frac{H_d}{H} = 3.748 - 10.977 \left(\frac{H}{H_o}\right) + 1.329 \left(\frac{S}{S_o}\right) + 7.667 \left(\frac{H}{H_o}\right)^2 - 1.044 \left(\frac{S}{S_o}\right)^3$	35
36	$\frac{H}{H_o}, \frac{S}{S_o}$	$\frac{H_d}{H} = 2.740 - 6.126 \left(\frac{H}{H_o}\right) + 1.301 \left(\frac{S}{S_o}\right) - 1.029 \left(\frac{S}{S_o}\right)^3 + 4.036 \left(\frac{H}{H_o}\right)^3$	36
37	$\frac{H}{H_o}, \frac{S}{S_o}$	$\frac{H_d}{H_o} = 0.557 - 0.136 \left(\frac{H}{H_o}\right) - 0.279 \left(\frac{S}{S_o}\right)$	37
38	$\frac{H}{H_o}, \frac{S}{S_o}$	$\frac{H_d}{H_o} = -0.627 + 3.758 \left(\frac{H}{H_o}\right) - 0.306 \left(\frac{S}{S_o}\right) - 3.139 \left(\frac{H}{H_o}\right)^2$	38
39	$\frac{H}{H_o}$, $\frac{S}{S_o}$	$\frac{H_d}{H_o} = 1.325 - 5.105 \left(\frac{H}{H_o}\right) + 2.125 \left(\frac{S}{S_o}\right) + 3.705 \left(\frac{H}{H_o}\right)^2 - 1.602 \left(\frac{S}{S_o}\right)^2$	39
40	$\exp\frac{H}{H_o}, \frac{S}{S_o}$	$\frac{H_d}{H} = 1.524 - 0.228 \left(\exp \frac{S}{S_o} \right) - 0.984 \left(\frac{H}{H_o} \right)$	40
41	$\frac{H}{H_o}$, exp $\frac{S}{S_o}$	$\frac{H_d}{H} = 1.260 - 0.230 \left(\exp \frac{S}{S_o} \right) - 0.111 \left(\frac{H}{H_o} \right) - 0.705 \left(\frac{H}{H_o} \right)^2$	41
42	$\frac{H}{H_o}$, exp $\frac{S}{S_o}$	$\frac{H_d}{H} = 3.459 + 0.886 \left(\exp \frac{S}{S_o} \right) - 9.41 \left(\frac{H}{H_o} \right) + 6.429 \left(\frac{H}{H_o} \right)^2 - 0.870 \left(\exp \frac{S}{S_o} \right)^2$	42
43	$\frac{H}{H_o}$, $In\frac{S}{S_o}$	$\frac{H_d}{H} = 0.880 - 0.890 \left(\frac{H}{H_o}\right) - 0.338 \left(In \frac{S}{S_o}\right)$	43

$$RRMSE = \frac{\left[\frac{1}{n}\sum_{i=1}^{n} (H_{d}^{est} - H_{d}^{mea})^{2}\right]^{\frac{1}{2}}}{\frac{1}{n}\sum_{i=1}^{n} H_{d}^{mea}} \times 100$$

$$R^{2} = 1 - \frac{\sum_{i=1}^{n} (H_{d}^{mea} - H_{d}^{est})^{2}}{\sum_{i=1}^{n} (H_{d}^{mea} - H_{d}^{avemea})^{2}}$$
(12)

$$R^{2} = 1 - \frac{\sum_{i=1}^{n} \left(H_{d}^{mea} - H_{d}^{est} \right)^{2}}{\sum_{i=1}^{n} \left(H_{d}^{mea} - H_{d}^{avemea} \right)^{2}}$$
(12)

where n in this section represents the total number of available data point (12), H_d^{est} = estimated diffuse solar radiation, H_d^{mea} = measured diffuse solar radiation, and H_d^{avemea} = average of measured diffuse solar radiation.

In order to check and select the best model out of the 43 recommended models used in this study, a global performance indicator (GPI) was applied. Applying the GPI newly introduced by Bailek et al. [37] and Despotovic et al. [51] on the 43 models developed in this study undoubtedly revealed the best model. The established best model was then used to compare the five measured data and five diffuse solar radiation models reported in the literature from different locations across the globe as presented in Tables 4 - 5. This was established in order to check if the accuracy and application of the best model is limited to the site from which the model was developed. This technique was applied because peers and researchers from the time immemorial reported that diffuse solar radiation is dependent on local climate and geographical location [22-24, 42, 50]. Also, GPI was applied in this section for selecting the best performing model out of 43 models and for comparing with the literature because of the following: (1) The GPI combines the advantages of the statistical indicators presented in equation (8 - 12) in order to reveal the best performing model, and (2) with the application of GPI, a single value which consists of short- and long-term statistical performances together with the linearity of the models will be clearly observed and selected. However, Bailek et al. [37] stressed that the GPI is a relative unbounded value and a higher value of the GPI implies a better statistical performance and modeling quality. According to Despotovic et al. [51] and Jamil and Akhtar [22], the values of all selected statistical indicators need to be scaled down so that the scaled values lie between 0 and 1. These scaled values are subtracted from the median value of the corresponding scaled statistical indicators. Finally, the values obtained are summed applying appropriate weight factors.

This indicator (GPI) is defined mathematically as:

$$GPI = \sum_{i=1}^{5} \alpha_j \left(\bar{y}_i - \bar{y}_{ij} \right)$$
 (13)

Where α_i equal -1 for the indicator R² only, whereas for other indicators (MBE, MPE, RMSE and RRMSE) is equal to 1. \bar{y}_i is the median of scaled values of the indicators j, \bar{y}_{ii} is the scaled value of indicator j for model i. The higher the values of GPI indicator, the better the accuracy of model.

Tal	ble 4.	Monthl	y mean dif	fuse solar	radiation	models	obtained	from 1	iterature
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Mo del	Parameters	Class	Regression relations	Reference
44	$\frac{H}{H_O}$	Clearness index-based model	$\frac{H_d}{H} = 1.0 - 1.13 \left(\frac{H}{H_o}\right)$	Page [25]
45	$\frac{H}{H_O}$	Clearness index-based model	$\frac{H_d}{H} = 0.958 - 0.982 \left(\frac{H}{H_o}\right)$	Iqbal [13]
46	$\frac{H}{H_o}$	Clearness index-based model	$\frac{H_d}{H} = 0.636 - 0.279 \left(\frac{H}{H_o}\right) - 0.194 \left(\frac{H}{H_o}\right)^2 - 0.383 \left(\frac{H}{H_o}\right)^3$	Ibraham [28]
47	$\frac{H}{H_O}$	Clearness index-based model	$\frac{H_d}{H} = 1.390 - 4.027 \left(\frac{H}{H_o}\right) + 5.531 \left(\frac{H}{H_o}\right)^2 - 3.108 \left(\frac{H}{H_o}\right)^3$	Liu & Jordan [49]
48	$\frac{H}{H_o}, \frac{S}{S_o}$	Clearness index and sunshine duration-based model	$\frac{H_d}{H_o} = 0.385 - 0.12 \left(\frac{H}{H_o}\right) + 0.067 \left(\frac{S}{S_o}\right)^2$	Maduekwe & Chendo [52]

Table 5. Measured monthly mean diffuse solar radiation obtained from the literature and the NASA website [62]

Lat.	Lon.	Ele.	Jan	Feb	Mar	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec
Keı	rman, Ir	an bety	veen 19	90-200	5 Safari	pour ar	d Mehr	abian [:	53]					
30.3 N	NA	NA	5.23	6.14	8.06	8.60	8.16	7.46	7.41	6.88	5.86	4.92	3.98	4.44
Ali	garh M	uslim U	Jniversi	ity, Alig	arh, Inc	lia betw	een 201	13-2016	5 Jamil	and Akl	ntar [22	.]		
27.9 N	78.1 E	NA	6.19	7.07	7.92	8.21	8.64	9.06	8.68	8.53	6.32	8.02	6.78	5.87
University of Tarapaca site of Arica-Parinacota, Chile between 2012-2014 Cornejo et al. [54]														
18.5 S	70.3 W	9 m	8.32	6.44	9.54	7.81	5.33	5.29	6.05	6.73	8.03	8.60	7.45	8.06
Las	Rejas,	Chile 2	2013-20	14 Cor	nejo <i>et d</i>	al. [54]								
18.3 S	69.5 W	4391m	8.32	7.06	8.35	3.60	3.38	3.35	7.67	3.13	3.13	5.33	5.04	8.93
NA	NASA, Hyderabad, Pakistan between 1983-2005 [55]													
65.4 N	68.3 E	60m	4.25	5.15	6.12	6.84	7.42	7.70	8.28	7.67	6.48	5.62	4.64	4.03

Where Lat. represents latitude positive north/south in degrees, Lon. stands for longitude positive east/west in degrees, Ele. denotes elevation in meters and monthly mean diffuse solar radiation obtained from literature and NASA (same geographical information as study site) are all in MJm⁻²day⁻¹

3. Results and Discussion

In this section, the results of the measured data in the study site were compared with the following: (1) the developed 43 models in this study, (2) five measured data obtained from the literature together with the observed satellite data obtained from the NASA database, and (3) five models obtained from the literature and best performing model (model 42) as presented in Figs. 2 - 5, and their corresponding estimation statistical indicators are presented in Tables 6 - 7.

3.1 Analysis of Monthly and Yearly Solar Radiation Ground Observation

The results of the monthly and annual averages of the study site and the corresponding aggregate mean values for the duration of measurement (2 years) are

presented. It can be seen vividly that the mean monthly and yearly solar radiation and meteorological values for the two years of measurement are presented in Table 2.

The summer season is from March-August whereas the winter season starts in October and ends in January for Hyderabad, Pakistan. The winter months are characterized by overcast, heavy rainfall clouds, heavy frogs, high relative humidity, low temperature, low wind speed and direction, and the highest ambient air pressure as shown in Table 2. This gives rise to minimum values of 14.33 MJm⁻²day⁻¹ and 14.75 MJm⁻²day⁻¹ reported for global and direct solar radiation in the month of January, respectively, whereas the minimum diffuse solar radiation of 5.02 MJm⁻²day⁻¹ was recorded in the month of December as shown in Table 2.

However, summer months are characterized by clear sky, high temperature and wind speed with low relative humidity and ambient air pressure. This culminates into high values of global solar radiation and its component. The maximum value for global solar radiation (24.84 MJm⁻²day⁻¹) occurred in the month of April, direct normal irradiance (20.78 MJm⁻²day⁻¹) occurred in the month of February, and diffuse solar radiation (13.28 MJm⁻²day⁻¹) occurred in the month of July. These results are comparable to the report of Jamil and Akhtar [22-24] in the humid-subtropical climatic region of India.

The yearly averages of direct normal irradiance, global solar radiation and diffuse solar radiation calculated for the period of measurement May 2015-April 2016 are 6430.141 MJm⁻²year⁻¹, and 7593.977 MJm⁻²year⁻¹, respectively. For the interval of May 2015-April 2017, the calculated values of global solar radiation, direct normal irradiance and diffuse solar radiation are 7503.759 MJm⁻²year⁻¹, and 3153.287 MJm⁻²year⁻¹, respectively. While for only year 2016 (January-December), the calculated values of global solar radiation, direct normal irradiance and diffuse solar radiation are 7500.64 MJm⁻²year⁻¹, 6350.196 MJm⁻²year⁻¹, and 3139.756 MJm⁻²year⁻¹, respectively. It could be observed that the yearly global solar radiation is higher than the yearly direct normal irradiance with about 15.2% between May 2015-April 2016, and 18.3% in year 2016 (January-December) as shown in Table 2. This implies that the site is more favorable for the installation of photovoltaic technology or flat solar collectors as the magnitude of direct normal irradiance is below the threshold of 7200 MJm⁻²year⁻¹ in the months of January, June, July, August, November and December as presented in Table 2. It is therefore imperative to note that concentrated solar power should not be considered as a favorable technology in this station.

Avoiding the energy available for the development of solar power technologies in Hyderabad, Pakistan, various other factors and technical aspects reduce the actual use of solar energy captured [54, 56-62].

In Hyderabad, the global solar radiation is significantly higher than the direct normal irradiance as a result of higher attenuation effect of aerosols and water vapor on direct irradiance than the diffuse solar radiation component [54, 62]. Hyderabad is situated by the coast and it is highly affected by a high load of sea salt, and water drops aerosols and water vapor loads as the station is located in the University setting where thousands of people and building structures are located. As a result, atmospheric particles are able to absorb light beams of a specific wavelength. These particles convert electromagnetic radiation into heat and eventually into diffused solar radiation components. So, the direct normal irradiance is obtained from the relation:

$$H_b = H\cos z - H_d \tag{14}$$

where z is the zenith angle and other variables retain their usual meaning.

From equation (14), it is obvious that as the diffuse solar radiation component increases, the direct normal irradiance decreases and finally the global solar radiation remains at the same level.

3.2 Performance Evaluation

Developed models under the three classes of clearness index-based models, sunshine duration-based models and combination of the two predictors using either diffuse fraction or diffuse coefficient are now evaluated and results are shown in Fig. 2.

In clearness index-based models' class, diffuse fraction and diffuse coefficient are developed with only one predictor of the clearness index. Fourteen models are developed, with the restriction of the order of two in each input predictor. This could be attributed to the fact that the higher order equations have increased complexity. The numerous models developed are presented in Table 3a. The statistical indicators such as MBE, MPE, RMSE, RRMSE and R² have been evaluated for the developed models in the class. The results of the statistical indicators are presented in Table 6. MBE values lie in the range of -0.01057 to -0.00994 MJm⁻²day⁻¹ with a minimum value observed for model 6 (-0.01057 MJm⁻²day⁻¹). As observed from Table 6, models 1-14 under this class recorded negative values. This implies that models underestimated the measured data. However, the overestimation in the values is significantly small because the magnitude of MBE for this class is observably close to zero. MPE values lie in the range -0.08687 to -0.04886 MJm⁻²day⁻¹ with the minimum value observed for model 13 (-0.08687 MJm⁻²day⁻¹). RMSE values are observed to be small for all the developed models under this class with model 6 registering the minimum value of 0.177211 MJm⁻²day⁻¹). In general, the RMSE range buried between 0.177211 to 0.194205 MJm⁻²day⁻¹. Accordingly, the RRMSE value buried in the range 25.50065 to 28.04482 MJm⁻²day⁻¹ with the minimum value 25.50065 registered for model 6. The coefficient of determination (R²) has values in the range of 0.931 - 0.937 representing good fit of measured data. The highest value of R² was recorded for model 5 and model 13.

Under the sunshine duration-based models, the diffuse fraction and diffusion coefficient models are developed with only one predictor of the sunshine duration parameter. Seventeen models are proposed, with the restriction of the order of two in each input predictor. This is because, the higher order equations demonstrate increased complexity. Hence, several models proposed under this class are presented in Table 3b. From the statistical indicators evaluated under this class, the results are presented in Table 4. MBE value lies in the range of -0.0043 to 0.003714 MJm⁻²day⁻¹ with the minimum value of -0.0043 MJm⁻²day⁻¹ registered for model 26. As observed models 15, 17, 22, 25, 28, 29 and 30 recorded a positive value of MBE, indicating overestimation while the remaining models reported a negative value implying an underestimation. However, the overestimation and underestimation in the values is significantly small since the values of MBE for the proposed models reasonably close to zero. This trend is equally observed for the models developed in the humid-subtropical climate region of India [22]. MPE values lie in the range -0.07466 to 0.005151 MJm⁻²day⁻¹ with the minimum value of -0.07466 MJm⁻²day⁻¹ recorded for model 29. RMSE values recorded small values for all the developed models under this class with the minimum value of 0.134126 MJm⁻²day⁻¹ registered for model 25, from range of values 0.13416 to 0.158558 MJm⁻²day⁻¹. Accordingly, the RRMSE value buried in the range 19.01901 to 23.50345 with the minimum value of 19.01901 reported for model 26. The coefficient of determination (R^2) has values in the range of 0.943 - 0.961 indicating good fit of the measured data. The highest value of R^2 is reported for models 20-22 and model 25-26.

Table 6. The statistical test indicators of all developed models for estimating the diffuse solar radiation at

Models MBE MPE RMSE RRMSE R2 GPI Rank	Hyderaba	d, Pakistan							
1	Models	MBE	MPE	RMSE	RRMSE	\mathbb{R}^2	GPI	Rank	
2 -0.00886 -0.06974	Clearness index-based models								
3			-0.07739		27.82833		-2.2996		
4	2								
5 -0.00874 -0.08011 0.188957 27.33110 0.937 -2.1183 38 6 -0.01057 -0.04886 0.177650 25.50065 0.931 -1.5319 14 8 -0.00746 -0.08654 0.1778283 25.96952 0.932 -1.6400 17 9 -0.00824 -0.08428 0.187491 27.18009 0.936 -2.0729 36 10 -0.00932 -0.0219 0.190773 27.49131 0.934 -2.2654 41 11 -0.00878 -0.08250 0.194205 28.04482 0.932 -2.2969 42 12 -0.00878 -0.07119 0.17981 25.97848 0.932 -2.17402 23 13 -0.00876 -0.08411 0.178343 25.95240 0.932 -1.5986 14 Sunshine duration-based models 15 0.00139 -0.05115 0.155966 23.25241 0.946 -2.1489 39 16 -0.00023 -0.04845 0.15827 </td <td></td> <td>-0.00892</td> <td>-0.06916</td> <td>0.179255</td> <td>25.89493</td> <td></td> <td>-1.7088</td> <td></td>		-0.00892	-0.06916	0.179255	25.89493		-1.7088		
6		-0.00820	-0.07808	0.179834	26.03971		-1.7190		
7	5								
8									
9		-0.00994	-0.05646	0.177650		0.931	-1.5319	14	
10			-0.08654	0.178283	25.96952		-1.6400		
11	9	-0.00824	-0.08428			0.936	-2.0729	36	
12	10	-0.00932	-0.07219	0.190773	27.49131	0.934	-2.2654		
13		-0.00878	-0.08250	0.194205	28.04482	0.932	-2.2969		
14		-0.00878	-0.07119	0.179881	25.97848	0.932	-1.7402		
Sunshine duration-based models 15	13	-0.00827	-0.08687	0.189859	27.51315	0.937	-2.1177	37	
15	14	-0.00767	-0.08411	0.178343	25.95240	0.932	-1.5986	14	
16 -0.00023 -0.04845 0.158277 23.40913 0.948 -1.9760 31 17 -0.00081 -0.06078 0.158309 23.46320 0.948 -1.9760 31 18 -0.00094 -0.03919 0.158589 23.38298 0.947 -1.9916 33 19 -0.00179 -0.03799 0.136359 19.24055 0.958 -1.3733 10 20 -0.00249 -0.01617 0.135120 19.06948 0.961 -1.8065 28 21 -0.00082 -0.04779 0.134612 19.16734 0.961 -1.8027 27 22 0.000452 -0.05150 0.134273 19.15748 0.961 -1.796 26 23 -0.00193 -0.04544 0.139213 19.66207 0.955 -1483 12 24 -0.00216 -0.02825 0.135648 19.13818 0.959 -1.3835 11 25 0.00196 -0.07009 0.134126 19.22668 0.96	Sunshine	duration-based	models						
17 -0.00081 -0.06078 0.158309 23.46320 0.948 -1.9760 31 18 -0.00094 -0.03919 0.158589 23.38298 0.947 -1.9916 33 19 -0.00179 -0.03799 0.136359 19.24055 0.958 -1.3733 10 20 -0.00249 -0.01617 0.135120 19.06948 0.961 -1.8065 28 21 -0.00082 -0.04779 0.134612 19.16734 0.961 -1.8027 27 22 0.000452 -0.05150 0.134273 19.15748 0.961 -1.796 26 23 -0.00193 -0.04544 0.139213 19.66207 0.955 -1483 12 24 -0.00216 -0.02825 0.135648 19.13818 0.9961 -1.7812 25 26 -0.0043 0.005151 0.135586 19.01901 0.961 -1.7549 24 27 -0.00121 -0.03128 0.157561 23.29592 0.9	15	0.00139	-0.05115	0.155966	23.25241	0.946	-2.1489	39	
18 -0.00094 -0.0319 0.158589 23.38298 0.947 -1.9916 33 19 -0.00179 -0.03799 0.136359 19.24055 0.958 -1.3733 10 20 -0.00249 -0.01617 0.135120 19.06948 0.961 -1.8065 28 21 -0.00082 -0.04779 0.134612 19.16734 0.961 -1.8027 27 22 0.000452 -0.05150 0.134273 19.15748 0.961 -1.796 26 23 -0.00193 -0.04544 0.139213 19.66207 0.955 -1483 12 24 -0.00216 -0.02825 0.135648 19.13818 0.959 -1.3835 11 25 0.00196 -0.07009 0.134126 19.22668 0.961 -1.7812 25 26 -0.0043 0.005151 0.135586 19.01901 0.961 -1.7549 24 27 -0.00121 -0.03128 0.157561 23.31739 0.948<	16	-0.00023	-0.04845	0.158277	23.40913	0.948	-1.9851	32	
19	17	-0.00081	-0.06078	0.158309	23.46320	0.948	-1.9760	31	
20 -0.00249 -0.01617 0.135120 19.06948 0.961 -1.8065 28 21 -0.00082 -0.04779 0.134612 19.16734 0.961 -1.8027 27 22 0.000452 -0.05150 0.134273 19.15748 0.961 -1.796 26 23 -0.00193 -0.04544 0.139213 19.66207 0.955 -1483 12 24 -0.00216 -0.02825 0.135648 19.13818 0.959 -1.3835 11 25 0.00196 -0.07009 0.134126 19.22668 0.961 -1.7812 25 26 -0.0043 0.005151 0.135586 19.01901 0.961 -1.7549 24 27 -0.00121 -0.03128 0.157561 23.29592 0.948 -1.52144 13 28 0.00079 -0.05283 0.15682 23.31739 0.948 -2.0489 35 29 0.003714 -0.07466 0.155199 23.31988 0.943<	18	-0.00094	-0.03919	0.158589	23.38298	0.947	-1.9916	33	
21 -0.00082 -0.04779 0.134612 19.16734 0.961 -1.8027 27 22 0.000452 -0.05150 0.134273 19.15748 0.961 -1.796 26 23 -0.00193 -0.04544 0.139213 19.66207 0.955 -1483 12 24 -0.00216 -0.02825 0.135648 19.13818 0.959 -1.3835 11 25 0.00196 -0.07009 0.134126 19.22668 0.961 -1.7812 25 26 -0.0043 0.005151 0.135586 19.01901 0.961 -1.7549 24 27 -0.00121 -0.03128 0.157561 23.29592 0.948 -1.52144 13 28 0.00079 -0.05283 0.15682 23.31739 0.948 -2.0489 35 29 0.003714 -0.07466 0.155199 23.31988 0.943 -2.2366 40 30 0.001417 -0.06818 0.15855 23.50345 0.947 </td <td>19</td> <td>-0.00179</td> <td>-0.03799</td> <td>0.136359</td> <td>19.24055</td> <td>0.958</td> <td>-1.3733</td> <td>10</td>	19	-0.00179	-0.03799	0.136359	19.24055	0.958	-1.3733	10	
22 0.000452 -0.05150 0.134273 19.15748 0.961 -1.796 26 23 -0.00193 -0.04544 0.139213 19.66207 0.955 -1483 12 24 -0.00216 -0.02825 0.135648 19.13818 0.959 -1.3835 11 25 0.00196 -0.07009 0.134126 19.22668 0.961 -1.7812 25 26 -0.0043 0.005151 0.135586 19.01901 0.961 -1.7549 24 27 -0.00121 -0.03128 0.157561 23.29592 0.948 -1.52144 13 28 0.00079 -0.05283 0.15682 23.31739 0.948 -2.0489 35 29 0.003714 -0.07466 0.155199 23.31988 0.943 -2.2366 40 30 0.001417 -0.06818 0.158558 23.50345 0.947 -1.9659 30 31 -0.00187 -0.02838 0.158044 23.30298 0.948<	20	-0.00249	-0.01617	0.135120	19.06948	0.961	-1.8065	28	
23 -0.00193 -0.04544 0.139213 19.66207 0.955 -1483 12 24 -0.00216 -0.02825 0.135648 19.13818 0.959 -1.3835 11 25 0.00196 -0.07009 0.134126 19.22668 0.961 -1.7812 25 26 -0.0043 0.005151 0.135586 19.01901 0.961 -1.7549 24 27 -0.00121 -0.03128 0.157561 23.29592 0.948 -1.52144 13 28 0.00079 -0.05283 0.15682 23.31739 0.948 -2.0489 35 29 0.003714 -0.07466 0.155199 23.31988 0.943 -2.2366 40 30 0.001417 -0.06818 0.158558 23.50345 0.947 -1.9659 30 31 -0.00187 -0.02838 0.158044 23.30298 0.948 -2.0032 34 Sunshine duration and clearness index-based models 32 -0.00191 -0.04025	21	-0.00082	-0.04779	0.134612	19.16734	0.961	-1.8027	27	
24 -0.00216 -0.02825 0.135648 19.13818 0.959 -1.3835 11 25 0.00196 -0.07009 0.134126 19.22668 0.961 -1.7812 25 26 -0.0043 0.005151 0.135586 19.01901 0.961 -1.7549 24 27 -0.00121 -0.03128 0.157561 23.29592 0.948 -1.52144 13 28 0.00079 -0.05283 0.15682 23.31739 0.948 -2.0489 35 29 0.003714 -0.07466 0.155199 23.31988 0.943 -2.2366 40 30 0.001417 -0.06818 0.158558 23.50345 0.947 -1.9659 30 31 -0.00187 -0.02838 0.158044 23.30298 0.948 -2.0032 34 Sunshine duration and clearness index-based models 32 -0.00191 -0.04025 0.130386 17.70534 0.960 -1.2243 6 33 -0.00279<	22	0.000452	-0.05150	0.134273	19.15748	0.961	-1.796	26	
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26 -0.0043 0.005151 0.135586 19.01901 0.961 -1.7549 24 27 -0.00121 -0.03128 0.157561 23.29592 0.948 -1.52144 13 28 0.00079 -0.05283 0.15682 23.31739 0.948 -2.0489 35 29 0.003714 -0.07466 0.155199 23.31988 0.943 -2.2366 40 30 0.001417 -0.06818 0.158558 23.50345 0.947 -1.9659 30 31 -0.00187 -0.02838 0.158044 23.30298 0.948 -2.0032 34 Sunshine duration and clearness index-based models 32 -0.00191 -0.04025 0.130386 17.70534 0.960 -1.2243 6 33 -0.00279 -0.02738 0.130494 17.57645 0.960 -1.250 7 34 -0.00188 -0.02298 0.110254 14.69178 0.967 -1.1409 4 35 -0.00129 <td>24</td> <td>-0.00216</td> <td>-0.02825</td> <td>0.135648</td> <td>19.13818</td> <td>0.959</td> <td>-1.3835</td> <td>11</td>	24	-0.00216	-0.02825	0.135648	19.13818	0.959	-1.3835	11	
27 -0.00121 -0.03128 0.157561 23.29592 0.948 -1.52144 13 28 0.00079 -0.05283 0.15682 23.31739 0.948 -2.0489 35 29 0.003714 -0.07466 0.155199 23.31988 0.943 -2.2366 40 30 0.001417 -0.06818 0.158558 23.50345 0.947 -1.9659 30 31 -0.00187 -0.02838 0.158044 23.30298 0.948 -2.0032 34 Sunshine duration and clearness index-based models 32 -0.00191 -0.04025 0.130386 17.70534 0.960 -1.2243 6 33 -0.00279 -0.02738 0.130494 17.57645 0.960 -1.243 6 34 -0.00188 -0.02298 0.110254 14.69178 0.967 -1.1179 3 36 -0.00129 -0.02979 0.109897 14.62919 0.967 -1.081 2 37 -0.00137 <td>25</td> <td>0.00196</td> <td>-0.07009</td> <td>0.134126</td> <td>19.22668</td> <td>0.961</td> <td>-1.7812</td> <td>25</td>	25	0.00196	-0.07009	0.134126	19.22668	0.961	-1.7812	25	
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29 0.003714 -0.07466 0.155199 23.31988 0.943 -2.2366 40 30 0.001417 -0.06818 0.158558 23.50345 0.947 -1.9659 30 31 -0.00187 -0.02838 0.158044 23.30298 0.948 -2.0032 34 Sunshine duration and clearness index-based models 32 -0.00191 -0.04025 0.130386 17.70534 0.960 -1.2243 6 33 -0.00279 -0.02738 0.130494 17.57645 0.960 -1.250 7 34 -0.00188 -0.02298 0.110254 14.69178 0.967 -1.1409 4 35 -0.00129 -0.02979 0.10975 14.67827 0.967 -1.1179 3 36 -0.00186 -0.02261 0.109897 14.62919 0.967 -1.081 2 37 -0.00137 -0.04969 0.133455 18.50554 0.957 -1.2583 8 38 -0.00401 -0.01166 0.133968 18.05627 0.958 -1.2925 9 <td>27</td> <td>-0.00121</td> <td>-0.03128</td> <td>0.157561</td> <td>23.29592</td> <td>0.948</td> <td>-1.52144</td> <td>13</td>	27	-0.00121	-0.03128	0.157561	23.29592	0.948	-1.52144	13	
30 0.001417 -0.06818 0.158558 23.50345 0.947 -1.9659 30 31 -0.00187 -0.02838 0.158044 23.30298 0.948 -2.0032 34 Sunshine duration and clearness index-based models 32 -0.00191 -0.04025 0.130386 17.70534 0.960 -1.2243 6 33 -0.00279 -0.02738 0.130494 17.57645 0.960 -1.250 7 34 -0.00188 -0.02298 0.110254 14.69178 0.967 -1.1409 4 35 -0.00129 -0.02979 0.10975 14.67827 0.967 -1.1179 3 36 -0.00186 -0.02261 0.109897 14.62919 0.967 -1.081 2 37 -0.00137 -0.04969 0.133455 18.50554 0.957 -1.2583 8 38 -0.00401 -0.01166 0.133968 18.05627 0.958 -1.2925 9 39 -0.00326	28	0.00079	-0.05283	0.15682	23.31739	0.948	-2.0489	35	
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Sunshine duration and clearness index-based models 32 -0.00191 -0.04025 0.130386 17.70534 0.960 -1.2243 6 33 -0.00279 -0.02738 0.130494 17.57645 0.960 -1.250 7 34 -0.00188 -0.02298 0.110254 14.69178 0.967 -1.1409 4 35 -0.00129 -0.02979 0.10975 14.67827 0.967 -1.1179 3 36 -0.00186 -0.02261 0.109897 14.62919 0.967 -1.081 2 37 -0.00137 -0.04969 0.133455 18.50554 0.957 -1.2583 8 38 -0.00401 -0.01166 0.133968 18.05627 0.958 -1.2925 9 39 -0.00326 -0.00551 0.110418 14.61601 0.967 -1.1671 5 40 -0.00306 -0.02152 0.128539 17.31672 0.961 -1.7221 22 41 -0.00130	30	0.001417	-0.06818	0.158558	23.50345	0.947	-1.9659	30	
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Under sunshine duration and clearness index-based models, diffuse fraction and diffuse coefficient models are proposed with two predictors. Twelve models are developed, with the restriction of the order of three in each input predictor. This is as a

result of the fact that, the higher order equations have increased complexity. The numerous models developed are presented in Table 3c. From the statistical results, MBE, MPE and RMSE reported similar trends in signaling but varying magnitude as in clearness index-based models. Similar trend was reported for models developed in India [21] indicating that models employing clearness index and those combining clearness index and sunshine duration exhibit similar diurnal fluctuation.

Table 7. Test the best models to the models and measured data obtained from the literature and the NASA database [55]

database [5	05]							
Model/Site	MBE	MPE	RMSE	RRMSE	\mathbb{R}^2	GPI	Rank	Reference
Models from	n the literatur	e						
Model 23	-0.00193	-0.04544	0.139213	19.66207	0.955	0.2208	1	Present study
Model 44	-0.25264	2.799933	0.875162	116.391	0.953	-10.40736	6	Liu & Jordan [49]
Model 45	-0.22267	2.474549	0.771353	102.8651	0.953	-9.488693	4	Page [25]
Model 46	-0.13464	1.402934	0.466401	58.31886	0.951	-6.066975	3	Iqbal [13]
Model 47	-0.22582	2.494301	0.782266	103.6861	0.954	-9.583763	5	Ibrahim [28]
Model 48	0.063076	-1.03147	0.360443	56.31961	0.891	-0.416833	2	Maduekwu & Chendo [61]
Measured da	ata from the l	iterature toget	her with the d	ata from the N	IASA datab	ase		
Model 23 (best model)	-0.00193	-0.04544	0.139213	19.66207	0.955	0.1338	1	Present study
UTA, Chile	-0.11018	0.445126	0.861782	114.3659	0.376	-3.93826	4	Cornejo <i>et</i> al. [54]
Aligarh, India	-0.08497	0.598581	0.459634	57.5258	0.711	-0.848897	2	Jamil & Akhtar
LR, Chile	-0.25168	2.062406	1.182753	153.1898	0.202	-4.854534	5	[22] Cornejo <i>et</i> <i>al.</i> [54]
Kerman, Iran	-0.18323	1.938777	0.63473	80.59346	0.624	-3.826472	3	Safaripour & Mehrabian
Hydamahad	0.20269	2 248062	0.705552	02 49750	0.020	5 957610	6	[53]
Hyderabad, Pakistan	-0.20368	2.248962	0.705552	93.48759	0.939	-5.857619	6	NASA Data [55]

Under this section, five (5) empirical models are often used by researchers in the literature, and four (4) measured data obtained equally from the literature and the NASA website were employed to check the applicability of the best models from the 45 proposed models. The results of the statistical indicators evaluated under this section are presented in Table 7. MBE values lie in the range of 0.25264 to 0.063076 on the five models obtained from the literature while the range of MBE on measured data obtained from satellite data from NASA website is -0.25168 to -0.08497. From the resulting matrices, only models 48 [52] have a positive value of MBE which indicates overestimation while the remaining models and measured data have a negative value leading to an underestimation. However, the overestimation and underestimation of these values is significantly small since the values of MBE for each of the proposed models are reasonably close to zero. MPE values lie in the range -1.03147 to 2.799933 and 0.445126 to 2.248962 with a minimum value of -1.03147 and 0.445126 for models from the literature (models 44-48) and measured data obtained from literature together with NASA data, respectively. Base on measured data obtained from the literature together with

NASA data, RMSE values are found to be minimum for model 48 and for Aligarh location India [22], respectively. Also, coefficient of determination (R²) recorded values in the range of 0.891 to 0.956 and 0.202 to 0.939 with maximum values of 0.956 reported for models 47 [28] and 0.939 for NASA data on models from the literature and measured data together with NASA data, respectively.

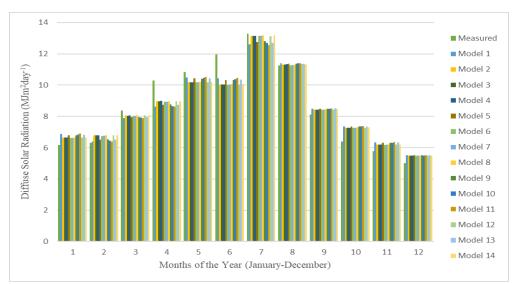


Fig. 2a. Clearness Index-based models for estimating diffuse solar radiation in Hyderabad, Pakistan

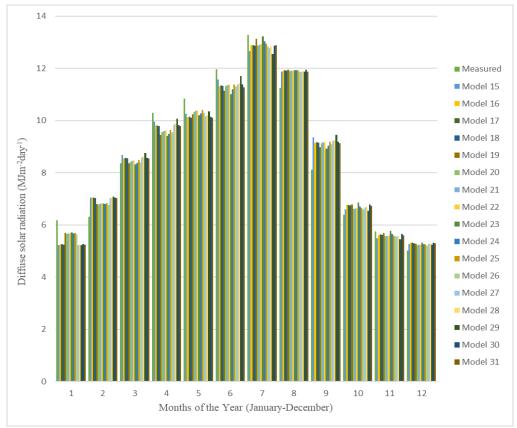


Fig. 2b. Sunshine duration-based models for estimating diffuse solar radiation in Hyderabad, Pakistan

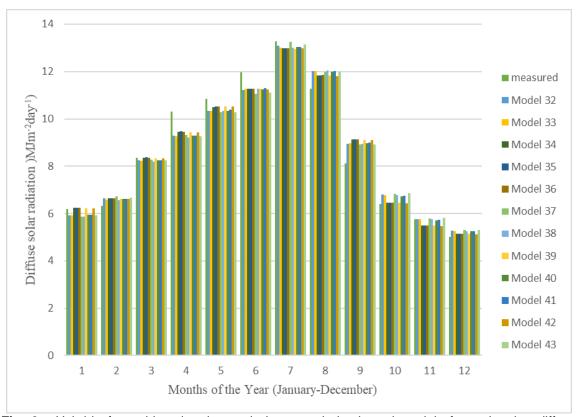


Fig. 2c. Hybrid of sunshine duration and clearness index-based models for estimating diffuse solar radiation in Hyderabad, Pakistan

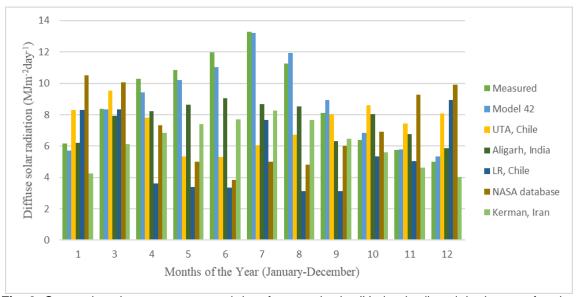


Fig. 3. Comparison between measured data from study site (Hyderabad) and the best performing model (model 42) together with measured the data obtained from the literature and the NASA database

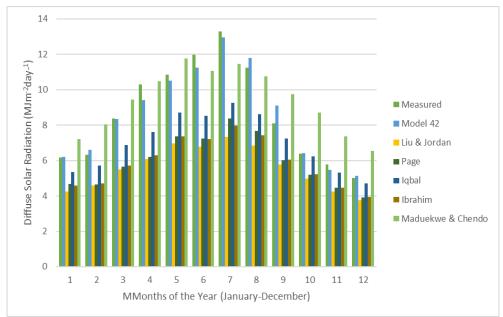


Fig. 4. Comparison between measured data from study site (Hyderabad) and the best performing model (model 42) together with models obtained from the literature

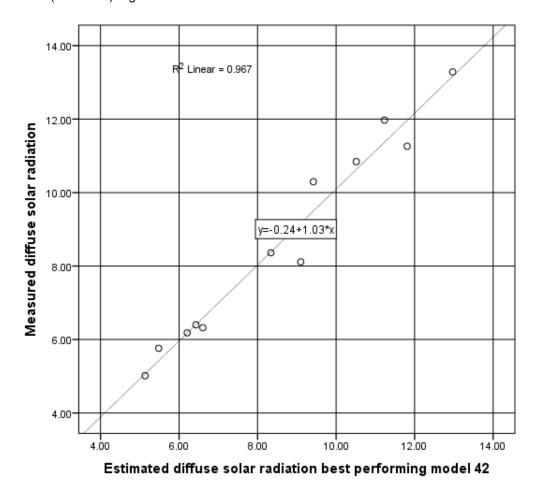


Fig. 5. Meanly mean variations of Hyderabad's best performing model after eliminating models from class 1-3 in $MJm^{-2}year^{-1}$

3.3 Global Performance Indicator and Ranking of Models

From the statistical indicators, it can be seen that different models from different classes come together with the models and measured data and satellite data obtained from the NASA database outperformed others. Thus, to avoid this variability and further improve the results of statistical analyses, global performance indicator (GPI) is applied.

As presented in Table 6, the GPI values of the proposed 43 models in this study classified under sunshine duration-based models, clearness index-based models and combination of both models are in the range of -2.2996 to -1.0208. The maximum GPI (-1.0208) and the minimum ranking of models were recorded for model 42 which is a hybrid of sunshine duration and clearness index predictors of variable dependent variable. It can be equally observed in Table 6 that hybrid of sunshine duration and clearness index predictors recorded as the best ranking model. This indicates that model 42 and hybrid of sunshine duration and clearness index predictors yielded the best performing model and class, respectively, in Hyderabad, Pakistan. Similar results were obtained in the literature [6, 20, 24, 41-42, 38-40].

In order to achieve the objective of the study, the best model (model 42) selected using GPI metric was applied to compare with five (5) models and five (5) measured data obtained from the literature and the NASA database. This is to check if the accuracy and applicability of the best model are limited from which the model was developed, as researchers and peers reported that diffuse solar radiation and other components of global solar radiation are dependent on local climate and regional geography.

As presented in Table 7, the maximum GPI and the minimum scores of the five (5) models and five (5) measured data from the literature together with satellite data obtained from the NASA database are compared with the best performing model in Hyderabad, Pakistan station (i.e., model 42). After thorough analysis using statistical indicators, GPI and ranking of models (Table 7), the best model (model 42) can be employed for estimating diffuse solar radiation in Kerman, Iran and Algarh, India while UTA and LR stations located in Chile require local calibration of model 42 to actually fit the measured data. In general, model 42 is best suited to fitt data at stations Aligarh, India followed by Kerman, Iran while other stations such as UTA and LR located in Chile need local calibration to actually fit the measured values in Hyderabad. However, it can be observed that model 42 did not fit the calculated data from the NASA database despite the fact that the data employed for modeling 42 and that of NASA possess the same geographical information. This could be attributed to the fact that NASA data is estimated under 20% error from existing models in the literature and different locations compared to the study site. Also, comparing models 42 with 5 models in the literature revealed that model 42 actually predicted values obtained from Iqbal's model [13] and Maduekwu & Chendo's models [52] while other models such as Liu & Jordan [49], Page [25], and Ibrahim [28] required local calibration to actually fit the values of model 42. On the whole, Iqbal's model is best suited to fit the values of the best performing model in Hyderabad (model 42), followed by Maduekwu & Chendo's model.

4. CONCLUSIONS

The knowledge of diffuse solar radiation is important for the design and development of solar system. In this study, forty-three models were analyzed, which

proposed correlations for the diffuse component of global solar radiation. The performance of these models was compared to measured diffuse solar radiation of Hyderabad station by applying various statistical indicators such as MBE, MPE, RMSE, RRMSE, R² and GPI. The diffuse fraction and diffuse coefficient were employed as the dependent variable to correlate with the sunshine duration fraction, clearness index and a combination of them. From the results of the statistical indicators, GPI and ranking of the models, model 42 - a hybrid of sunshine duration and clearness index-based models employing diffuse fraction dependent variable outperformed the remaining models proposed in this study. Moreover, the performance of the best model (model 42) was calculated equally with respect to a set of five (5) models and five (5) measured data of diffuse solar radiation presented in the literature together with satellite data obtained from the NASA database. The developed best model (model 42) in this study recorded favorable accurate results in comparison to these existing models and measured data of diffuse solar radiation in the literature and the NASA database with the lowest statistical indicators. The models were then categorized into rankings in descending GPI values so as to ascertain the best performer. From the results, model 42 recorded the highest value of GPI and was ranked first. Hence, it is concluded that the hybrid of sunshine duration and clearness index-based models of diffuse fraction:

$$\left(\frac{H_d}{H} = 3.459 + 0.886 \left(\exp\frac{S}{S_o}\right) - 9.411 \left(\frac{H}{H_o}\right) + 6.429 \left(\frac{H}{H_o}\right)^2 - 0.870 \left(\exp\frac{S}{S_o}\right)^2\right)$$

is selected as the most accurate generalized model for estimating solar radiation on the horizontal surface in Hyderabad of Pakistan, other locations such as Aligarh in India, Kerman in Iran, Iqbal's model and Maduekwu and Chendo's model, and in any location with similar climate conditions, in the absence of measured data.

This generalized model can be employed by solar energy engineers in terms of site selection and techno-economic performance quantification of solar energy applications such as photovoltaic technologies.

Nomenclature

$kt=H/H_o$	Monthly average clearness index (dimensionless)
H	Global solar radiation on the horizontal surface (MJm ⁻² day ⁻¹)
H_{o}	Monthly extraterrestrial solar radiation on the horizontal surface (MJm ⁻
	2 day $^{-1}$)
S_{o}	Maximum sunshine duration (hrs)
n	Number of days of the year
Isc	Solar constant in W/m ²
H_d	Monthly mean diffuse solar radiation (MJm ⁻² day ⁻¹)
Tave	Mean monthly maximum temperature (°C)
P	monthly mean atmospheric pressure at the site (hPa)
P_{o}	the standard atmospheric pressure at the sea level (1013 hPa)
S	Monthly mean sunshine hours (hrs)
S/S_o	sunshine fraction (dimensionless)
T_d	Monthly average dew point temperature
RH	Monthly mean relative humidity (%)

Greek Letters

- *φ* latitude (degrees)
- δ solar declination angle (degrees)
- ω sunset hour angle (degrees)

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CONFLICTS OF INTEREST

The author declares that there is no conflict of interests regarding the publication of this paper.

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